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GEODYNAMIC SERIES

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TABLE 1. Orbital Characteristics of Significant Satellites Used for Modeling the Earth's Gravity Field					
Satellite	Inclination, degrees	Altitude Range, km	Launch Date	Instrument	Approximate Accuracy
Agena Rocket Body	60.01	0.12-926	1964	O	3-4"
Anna-B	50.1	1071-1182	Oct. 1962	RR	3-4"
BE-B	70.69	876-1075	1964	L, RR,	5 cm/s
BE-C	41.19	980-1925	1965	O	2 m, 5 cm/s
DI-C	28.31	971-1211	1960	L, RR, O	3-4"
DI-D	30.97	578-1959	1967	O	3-4"
Echo-1	30.46	598-1890	1967	L, O	3-4"
GEOS-1	47.21	1494-16N2	1960	O	3-4"
GEOS-1	59.4	1116-2277	Nov. 6, 1965	L, RR, O	1-2 in, 5 cm/s
GEOS-2	105.8	671-976	Jan. 11, 1968	L, R, RR, O	1-2 m, 1 m
GEOS-3	115.0	821-854	April 10, 1975	L	5 cm/s
GRS	49.76	428-1294	1963	O	25 cm
INJUN	66.82	880-998	1961	O	3-4"
Landstar-1	98.8	898-911	July 29, 1972	RR	5 cm/sec
Mida-1	100.8	5898-5945	May 4, 1976	O	5 cm
OICO-2	95.83	3505-3729	1961	O	3-4"
OSCAR-7	87	413-1510	Oct. 1965	O	3-4"
OWI-2	89.7	867-1109	1966	O	3-4"
SECOR-5	144.27	1786-2092	1965	O	3-4"
Starkie	69.2	1195-2419	Aug. 10, 1965	O	3-4"
POLE-1	49.8	806-1108	Feb. 6, 1973	L, M	5 cm
POLE-1	15.0	395-465	Dec. 12, 1970	L, M	1 m
					20"

M: Minitrack; L: Laser Ranging; R: Radar Range; RR: Doppler; and O: Optical.

TABLE 2. Satellites That Have Measured the Near-Earth Geomagnetic Field

Satellite	Inclination, degrees	Altitude Range, km	Launch Dates	Instrument	Approximate Accuracy, nT
Sputnik 3	65	440-600	May-June 1958	Fluxgate	100
Vanguard 3	33	510-3780	Sept.-Dec. 1959	Proton	10
Vanguard 3	33	510-3780	Sept. 1965-Jan. 1974	Fluxgate	50-85
Cosmos 26	49	270-403	March 1964	Proton	Unknown
Cosmos 49	50	261-488	Oct.-Nov. 1964	Proton	22
OCO-8C	90	1040-1080	Dec. 1964-June 1966	Rubidium	22
OCO-4	87	413-1510	Oct. 1965	Rubidium	6
OCO-6	86	412-908	July 1967-Jan. 1969	Rubidium	6
Cosmos 921	82	997-1098	June 1969-July 1971	Rubidium	6
Auror	72	270-403	Jan.-March 1970	Cesium	Unknown
Auror	103	384-3145	Nov. 1969-June 1970	Fluxgate	Unknown
Titan	Polar	750-892	Sept. 1972-present	Fluxgate	Unknown
Magat	Polar	352-561	Oct. 1972-June 1980	Fluxgate	2
				Cesium	2

From Langel (1980).

GRM: Observing the Terrestrial Gravity and Magnetic Fields in the 1990's

P.T. Taylor, T. Keating,
W.D. Kahn, R.A. Langel,
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Beginning with the earliest days of space exploration, satellites have been used to chart the gravity and magnetic fields of the earth. As a continuation of these studies NASA is proposing to launch a new geopotential fields exploration system called the Geopotential Research Mission (GRM). Two spacecraft will be placed in a circular polar orbit at 100 km altitude. Distances between these satellites will vary from 100 to 600 km. Both scalar and vector magnetic fields will be measured by magnetometers mounted on a boom positioned in the forward direction on the lead satellite. Gravity data will be computed from the measured change in distance between the two spacecraft. This quantity, called the range-rate, will be determined from the varying frequency (Doppler shift) between transmitter and receiver on each satellite. Expected accuracies (at the one sigma level) are: gravity field, 1×10^{-3} in s^{-2} (1 milGal); 5 m in geoid height; magnetic scalar field 100 nT; vector to 20 arcseconds (96 microradians); both resolved to less than 100 km.

With these more accurate and higher resolution data we will be able to investigate the earth's structure from the crust (with the shorter wavelength gravity and magnetic anomalies) through the mantle (from the intermediate wavelength gravity field) and into the core (using the longer wavelength gravity and magnetic fields).

Introduction

From the very beginning of the artificial satellite era, space platforms have been used to map the gravity and magnetic fields of the earth. In 1958 Vanguard 1 (O'Keefe et al., 1959) and Sputnik 3 were the first satellites to map these geopotential fields. With the current field models (Lerch et al., 1981), we have a 500 km horizontal resolution; however, the GRM should provide a 100 km wavelength resolution.

Earth orbiting satellites are advantageous for geopotential field determinations owing to their rapid, global coverage. With magnetic

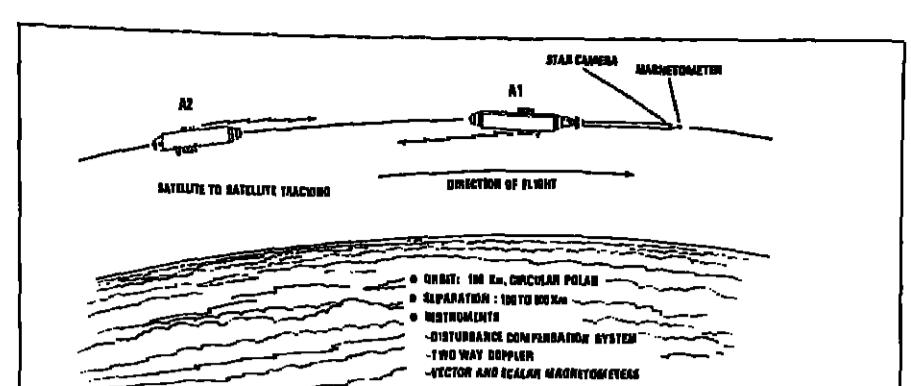
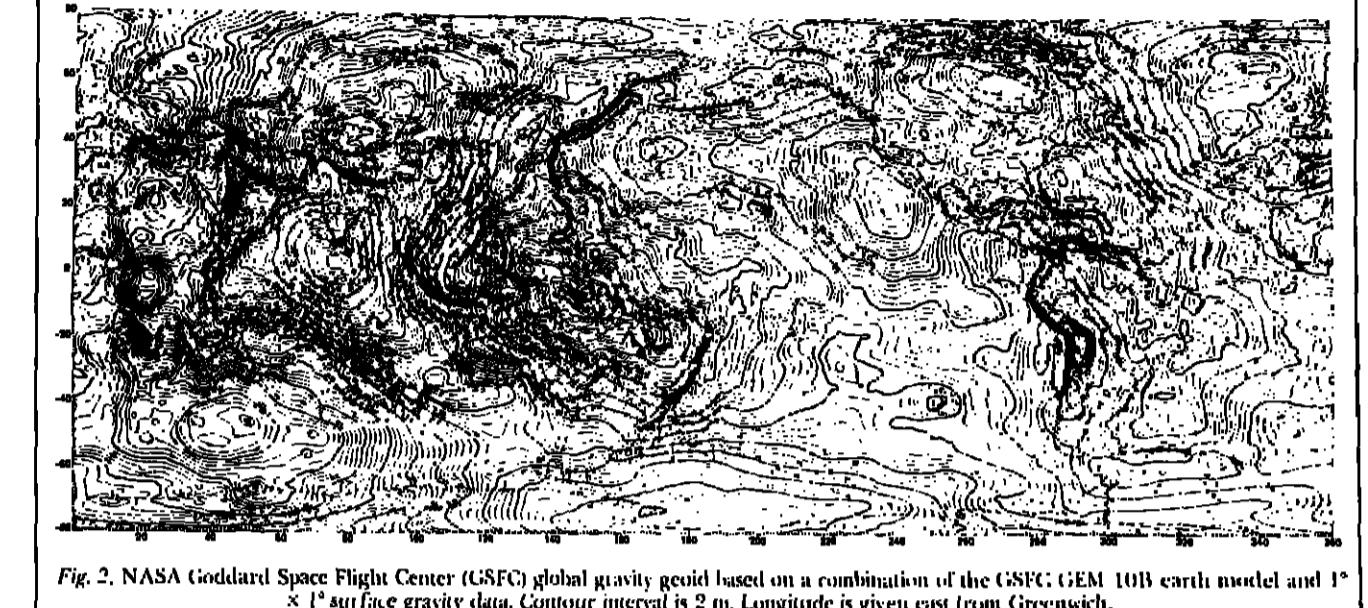


Fig. 1. GRM flight profile. Schematic representation of mission satellite configuration.



measurements, speed is particularly important since the field changes in time as well as space. These earlier missions have had two unsatisfactory characteristics: they were relatively high (>400 km) and had elliptical orbits. The Geopotential Research Mission will record both the gravity and magnetic fields from a 100 km altitude circular orbit. This project will unify the measurement of these major geopotential fields by recording them simultaneously. We will discuss this planned project and its scientific rationale and present some simulations of the anticipated data set.

To ensure complete global coverage, GRM will consist of two satellites in a $90^\circ \pm 0.1^\circ$ polar orbit of 100 km altitude for 6 months and be separated by up to 600 km (Figure 1). The operation altitude of 100 km was chosen to reconcile the scientific requirements for resolution with the engineering design for the size of the fuel tanks. These parameters are firm mission requirements and these spacecraft will be designed, for fuel and drag considerations, to fully meet these parameters. Gravity values will be derived from recording the range-rate of the two spacecrafts, while the lead vehicle will position a boom containing scalar and vector magnetometers and the attitude-determining sun sensors and star cameras.

The time of launch of this proposed new mission, if approved, will most likely be 1989-1990. To ensure complete global coverage, GRM will be launched into a 100 km circular-polar orbit, will measure both the gravity and magnetic fields of the earth with an accuracy and resolution not as yet obtainable from previous spacecraft. Unlike earlier satellite altimetry missions, GRM will record gravity field data over continental as well as oceanic areas. Having this accurate and detailed data set will allow us to study many regional scientific problems of the solid earth.

Both gravity and magnetic reference field models (such as in Figures 2 and 3) will be greatly improved by the data from GRM. A more detailed gravity field model will allow additional refinements to be made for the determination of other satellite orbits (for example, of altimeter bearing satellites). Previously completed altimetry missions (e.g., Seasat and GEOS 3) could have their orbital parameters redetermined yielding a much more accurate estimation of sea-surface height. Together with the vastly improved geoid, the redetermination will render the sea-surface topography in unprecedented detail.

Further improvements to the global geodetic datum will also be achieved. Since all geophysical anomaly fields are produced by removing the theoretical or reference field from the observed data, any significant improvement will result in refined anomaly mapping, particularly in the area of crustal studies.

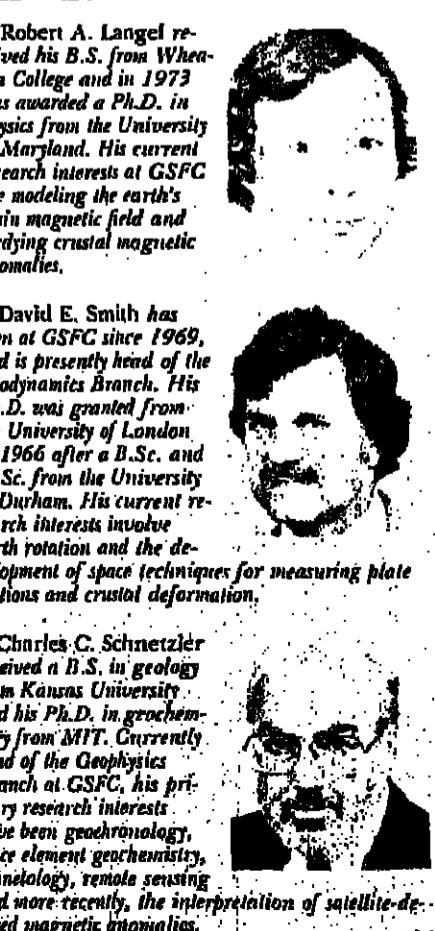
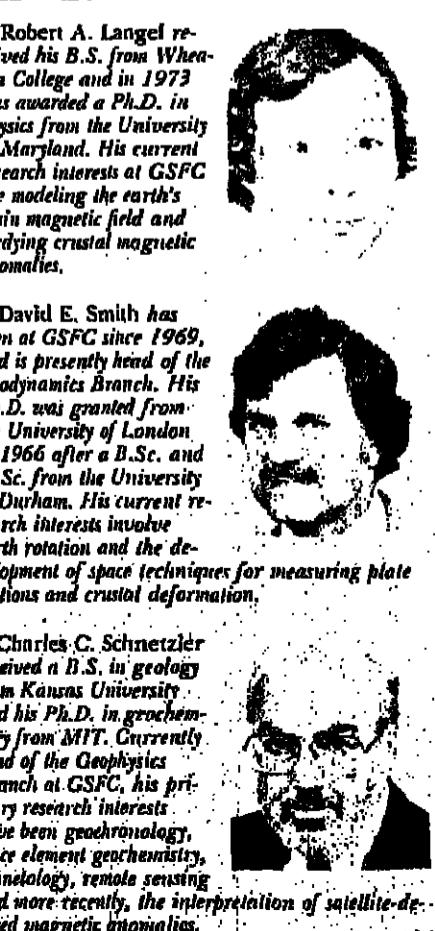
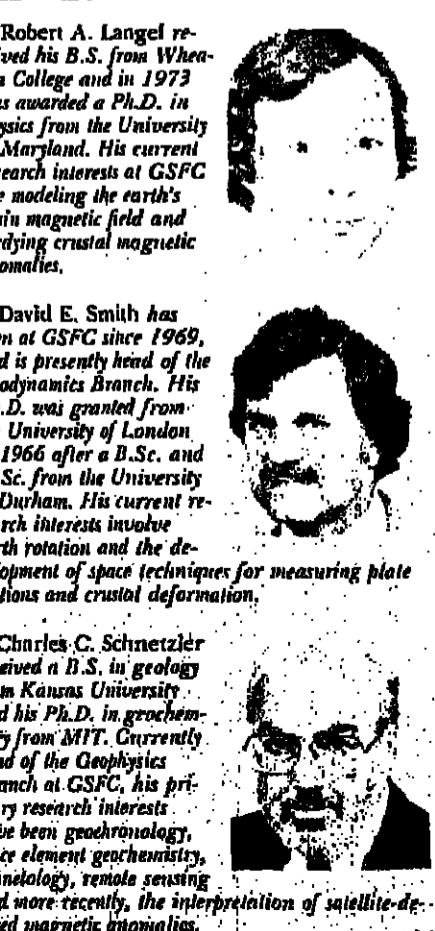
Enhancement of the intermediate wavelength features of the gravity field models ($100 \text{ km} \leq \lambda \leq 1000 \text{ km}$) are important for our understanding of the mass variation in the upper mantle. These mass anomalies are probably related to the forces driving the lithospheric plates as well as indicating the physical

Article (cont. on p. 610)

Robert A. Langel received his B.S. from Wheaton College and in 1973 was awarded a Ph.D. in physics from the University of Maryland. His current research interests at GSFC are modeling the earth's main magnetic field and studying crustal magnetic anomalies.

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cal characteristics of material in the asthenosphere. Shorter wavelength features ($\lambda < 100$ km) are indicative of the relative density, nature, and physical state of the lithosphere and crust. Marsh and Marsh [1976] and Louck [1976] have correlated patterns in the global free-air gravity anomaly field with geological features, convection, and density inhomogeneities.

Magnetic field data obtained from GRM will enable comparable insights into our understanding of the solid earth. Both the vector and scalar magnetic field measurements will be used to further refine the geomagnetic reference field models, which are being continually improved (Figure 3). When the dense and precise GRM values are used alone or combined with the earlier reference field [Langel et al., 1980] we will sense the time-varying magnetic field as never before. The time terms derived from this comparison can be applied to the commonly used International Geomagnetic Reference Field [Pettie, 1981] for the practical applications of magnetic survey reductions and compilation of areally contiguous magnetic charts. In maintaining global reference field models it is necessary to record the magnetic field periodically in order to detect the unsteady temporal field variations.

Beyond the mechanics of refining the reference fields for chart reduction and navigational purposes, a great deal can be learned about the earth's interior from these studies. The energy distribution in the lower harmonics in the magnetic field are important parameters in studies of the earth's core [Bentley et al., 1979].

Shorter wavelength features ($\lambda \leq 500$ km) of the magnetic field, the so-called anomaly field, have been interpreted to be the result of: (1) intracrustal contrasts in magnetization [Regan and Marsh, 1982; Mayhew et al., 1982]; (2) variations in the thickness of the crustal magnetized layer by Curie isotherm depth [Mayhew, 1982] or crustal thickness [Schneidler and Allende, 1983]; and (3) surface geological/tecnologic features [Frey, 1982].

With the greater increase in resolution of the magnetic anomaly field from the GRM it is possible to derive a geological/technological framework for economic evaluation as well as solve significant plate-tectonic problems.

The gravity and magnetic field data sets obtained by GRM will be complementary and supplementary. Both will furnish information about the lithosphere, its bulk composition, physical properties, and state. But gravity data will be primarily sensitive to mantle inhomogeneities and processes while magnetic measurements will mainly reflect the nature of the core. These geophysical potential field results will truly enable us to study the entire earth. These data sets will be useful to the studies of the smaller scale ($\lambda \approx 100$ km) global-crustal features such as foldbelts, rifts, structural basins, and oceanic fracture zones and seamounts.

Mission Profile

Two satellites (Figure 1) designated A1 and A2 will be launched by the shuttle into a polar ($90^\circ \pm 0.1^\circ$) orbit. The shuttle will launch the satellites at 275 km altitude, and they will self-deboost to a lower operating altitude of 160 km to begin a 6-month scientific mission. The spacecraft will orbit the earth with a period of about 88.5 minutes, providing approximately 16.3 orbits per day. Precision tracking of both spacecraft will be determined by the Defense Mapping Agency satellite tracking network.

Magnetometers, star cameras, and sun sensors will be located on an aluminum boom projected about 4 m from the A1 spacecraft (Figure 1). This lead satellite will weigh about 2800 kg, which is 200 kg greater than A2.

Gravity values will be derived by measuring the Doppler or range rate between the two satellites. The Satellite-to-Satellite Doppler Tracking System, which consists of ultrastable oscillators, multipliers, and horn antennas (the latter being mounted above and below each spacecraft (Figure 1)) will be able to detect small changes in the distance by measuring the Doppler frequency shift of the continuous wave of the 91 and 42 GHz signals.

According to the mission plan, the maximum distance between profile crossings of the equator will be about 7 km. The design goals for the earth's gravity field analysis are a 1 milligal accuracy in field measurement with a 5 cm geoid height resolution; both will be resolved to less than a degree (100 km).

These continuous wave signals are transmitted between the A1 and the A2 spacecraft. The Doppler-shifted signals will be compared to on-board reference frequencies, and consequently the variation in the range between the spacecraft will be determined. The range variation is continually time-sampled. The time-sampled range variation provides a measure of the relative velocity between the spacecraft to a sensitivity of 10^{-6} m s $^{-1}$.

Separating gravity field perturbations on the satellite orbits from external influences (such as atmospheric drag, solar radiation pressure, and solar and lunar gravity) is accomplished by the Disturbance Compensation System [Roy, 1983]. This system consists of a 14 cm diameter ball (called the "proof mass") within a 16 cm diameter spherical cavity. The spherical cavity is also an electrical capacitor. The position of the proof mass in the cavity is determined by the capacity relative to each axis. When the proof mass is off center, the resultant imbalance on the capacity is resolved into vector thrust commands to the appropriate thrusters of the propulsion system. The resultant movement of the spacecraft re-centers the proof mass within the cavity.

Since the spacecraft shields the proof mass from all perturbing surface forces, the mass responds only to variations in the gravity field. The two spacecraft positioned by the proof mass's response to gravity are themselves the instruments for detecting the gravity field. Interaction between the spacecraft and the earth's electromagnetic field could result in small, torque-like forces. These relatively weak electromagnetic field forces are countered by the reaction-wheel torque, thus eliminating any secular effects.

Distance between the spacecraft may be varied up to 600 km in order to more accurately measure specific harmonics of the gravity field spectrum.

According to the mission plan, the maximum distance between profile crossings of the equator will be about 7 km. The design goals for the earth's gravity field analysis are a 1 milligal accuracy in field measurement with a 5 cm geoid height resolution; both will be resolved to less than a degree (100 km).

Magnetic field determination should have the same spatial resolution with scalar field accuracy of 2 nT and vector components to 20 nT seconds. All stated accuracies are at the 1 s $^{-1}$ grid interval and contoured (Figure 4). Profiles were taken through this data and one is presented in Figure 5.

When we compare the upward-continued field (Figure 4) with a recent U.S. MagSat anomaly map at 320 km altitude (Figure 6) [Mayhew and Gaffiner, 1982], we can, not surprisingly, note the greater resolution of the lower orbiting satellite. Theoretical modeling studies, however, indicate that an even greater resolution of the magnetic anomaly field should be apparent. Differences in resolving power between the idealized models and the upward-continued U.S. magnetic anomaly field may be a result of a less than optimum field being removed from the latter. Further efforts are under way to define a more appropriate regional field for the digitized data base for the United States.

For gravity field simulations, a free-air anomaly map was computed at 160 km altitude of the new U.S. Geological Survey data base (Figure 7) [U.S. Geological Survey, 1982] using the same upward-continuation algorithm used with the magnetic field simulation. More efficient computer usage required that the original 4 km spacing between data points be increased to 20 km. Since we were computing the field at an altitude of 8 times this data interval, this was not considered to be a serious problem. These gravity values were plotted at the GRM nominal altitude (160 km) and at the nominal MagSat height (320 km). This comparison was used to suggest the increasing anomaly resolution to be expected from the lower-orbiting GRM.

In an effort to assess the geological significance of the potential field data anticipated from this proposed mission, a simplified tecnic map was superimposed on the magnetic field simulated at 160 km altitude (Figure 8). Since the magnetic measurements were not reduced to the pole, a strict spatial correlation cannot be made.

have revealed that the mission goals (gravity field to 2 milligals, geoid to 10 cm, and magnetic field anomaly map to 1 nT with a 20 arc-second direction reference—all with 100 km resolution) are realistic and should be achievable.

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News

Marine Sulfur Sources

It is of detecting and mapping lightning activity using many solid-state sensors because these sensors are most sensitive in the red rather than in the blue spectral region.

To test the hypothesis of spectral limitation, the lightning facility at the NASA Langley Research Center was used to produce an electrical discharge across a 5-mm gap in a special gas mixture to simulate lightning in the Venusian atmosphere. Both aluminum and tantalum electrodes were used to assess the continuation of electrode radiation.

A comparison of the air spectrum from the capacitor discharge with that obtained by Orville and L. E. Salama for terrestrial lightning showed that all the prominent lines in the terrestrial spectrum are also found in the spectrum from the capacitor discharge. Because the same lines appear with approximately the same relative intensity in both spectra, the capacitor discharge produces a useful simulation of terrestrial lightning; nevertheless, such a simulation must be regarded as probably exploratory.

A comparison of the spectra obtained from the simulation with those obtained by Orville and Salama for terrestrial lightning showed that (1) nitrogen lines are dominant in the terrestrial spectrum, but are not found in the simulated Venusian spectrum; (2) carbon lines are prominent in the simulated Venusian spectrum, but do not appear in the terrestrial spectrum; (3) oxygen lines and the hydrogen-H α line are present in both spectra.

A comparison of simulated and predicted spectra shows that most of the lines are present in the experimental spectrum. A comparison of the predicted radial fluxes from line radiation in the blue region (i.e., 350-800 nm) and the red region (600-900 nm) shows that the flux in the red region is approximately 5 times larger than that in the blue region.

It is fortunate that Venusian lightning is expected to radiate strongly in the region of 800 to 900 nm because the in situ measurements by V. I. Moroz et al. of atmospheric transmission, and the radiative transfer calculations by M. A. Williams et al., show that very little radiation at wavelengths shorter than 500 nm will penetrate the Venusian atmosphere and clouds. In particular, lightning observed through the Venusian atmosphere will have a reddish hue. Consequently, the reason the star sensor did not detect the Venusian lightning was not its spectral characteristics; rather, it seems likely that the low flashing rate or the low intensity produced a signal level that was too low to be detected.

This summary of a paper in the current *Geophysical Research Letters* was contributed by Joel D. Cline and B. Bates.

This summary of a paper in the current *Geophysical Research Letters* was contributed by William J. Borucki.

Geophysicists

Rear Admiral John D. Bostick, a member of the National Oceanic and Atmospheric Administration (NOAA) Corps and secretary of the AGU Geodesy section, is the new director of the National Oceanic and Atmospheric Administration's charting and geodetic services. Bostick was director of NOAA's National Geodetic Survey.

Frank Stehlé, dean of geosciences at the University of Oklahoma, has been appointed chairman of the Continental Scientific Drilling Committee of the Board on Earth Sciences, National Research Council.

Recent Ph.D.'s

For periodically lists information on recently accepted doctoral dissertations in the disciplines of geophysics. Faculty members are invited to submit the following information, on institution letterhead, above the signature of the faculty advisor or department chair: name of the dissertation title, author's name, name of the degree-granting department and institution, month and year degree was awarded. If possible include the current address and telephone number of the degree recipient (this information will not be published).

The Border Ranges Magmatic and Ultramafic Complex: Plutonic Core of an Intracanadian Volcanic Island Arc, Laurel E. Burns, Dept. of Geophysics, School of Earth Sciences, Stanford Univ., September 1983.

Crustal Shear Velocity Modeling in Nevada: A Study of Broadband Multi-Mode Surface Waves, Kin Yip Chun, Dept. of Geology and Geophysics, Univ. of California, Berkeley, December 1983.

Helium, Neutron, Water, and Carbon in Volcanic Rocks and Gases, Robert Joseph Poreda, Graduate Dept. of Scripps Institution of Oceanography, Univ. of California, San Diego, December 1983.

Venusian Lightning

Laboratory simulations by William J. Borucki and coworkers (*Geophysical Research Letters*, October 1983) indicate that the inability of satellite instruments to detect optical pulses of Venusian lightning is due to the weakness of Venusian lightning rather than its spectral characteristics.

Evidence of lightning activity in the Venusian atmosphere has been obtained by instrumentation aboard the Pioneer-Venus Orbiter (PVO) and Venus 9, 11, and 12 spacecraft. However, a search by Borucki and coworkers for optical pulses radiated by Venusian lightning was unsuccessful. The search used the geostationary sensor aboard the PVO to cover the geographical region from 70°N to 50°S latitude and from 240° to 60° longitude (IAU coordinates). B. Vonnegut and R. E. Orville have suggested that Venusian lightning might emit evidence of ionospheric discharges of greater than 400 nm. This would reduce the possibility of candidate experiments that would not

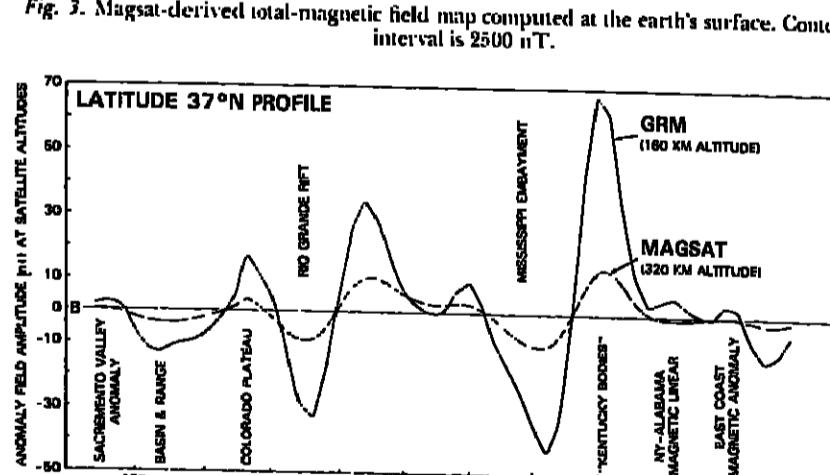
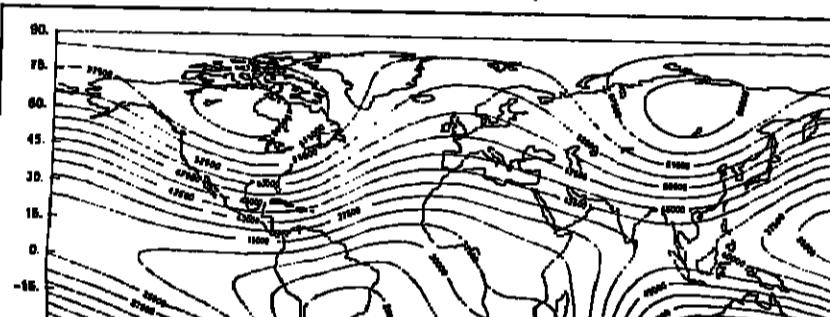


Fig. 5. Magnetic-anomaly field profiles along 37° north latitude. Data from Figures 4 and 6. Prominent tectonic/geophysical features are given for reference.

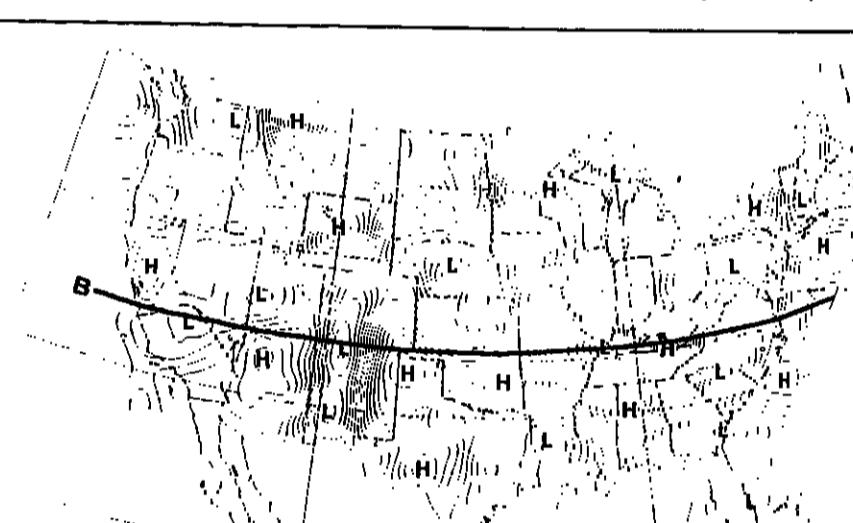


Fig. 4. U.S. Composite Magnetic Anomaly Map continued upward to 160 km altitude. Contour interval is 5 nT.

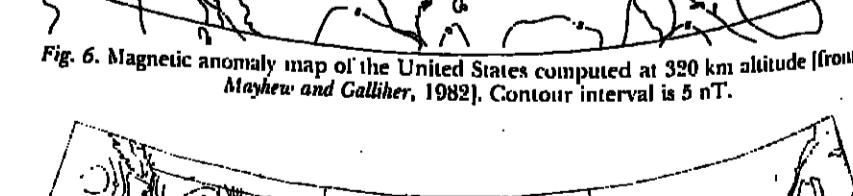


Fig. 6. Magnetic anomaly map of the United States computed at 320 km altitude [from Mayhew and Gaffiner, 1982]. Contour interval is 5 nT.

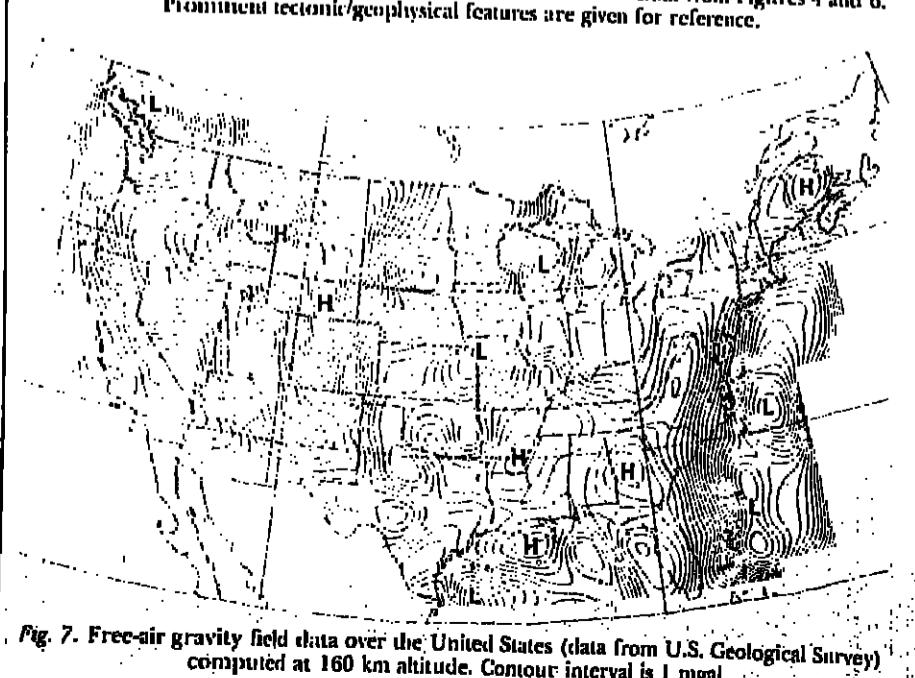


Fig. 7. Free-air gravity field data over the United States (data from U.S. Geological Survey) computed at 160 km altitude. Contour interval is 1 mgal.

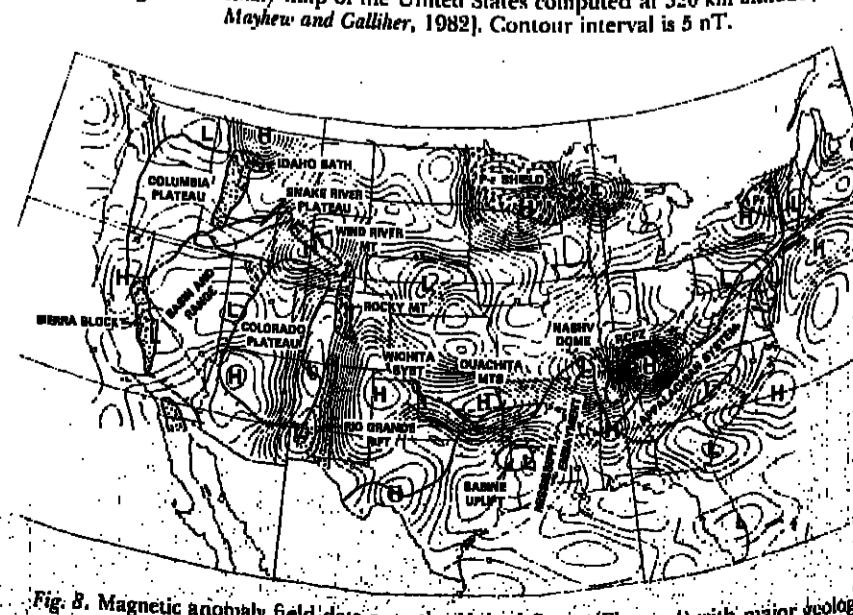


Fig. 8. Magnetic anomaly field data over the United States (Figure 4) with major geological features superimposed (based on King, 1977).



Section Candidates

AGU is carrying biographies and photographs of all candidates for President-elect, General Secretary, and Foreign Secretary of the Union and for President-elect and Secretary of each Section. In addition, statements by the candidates for Union offices and for Section President-elect will appear. The material for the Hydrology Section appears below. The material for the sections of Geodesy, Geomagnetism and Paleomagnetism, and Planetary appeared in the August 30 issue; for the Atmospheric Sciences Section in the September 27 issue; for the Tectonophysics section in the October 11; and for the Seismology Section in the October 18 issue. The slate of candidates for all offices was carried in the June 21 issue.

Hydrology: President-elect

Jacques W. Delleur A member of AGU since 1957; 58 years old; Professor of Hydrology and Hydraulics, Head of the Hydraulic and Systems Engineering Area of the School of Civil Engineering, Acting Director of Water Resources Research Center, Purdue Univ. Major interests: stochastic hydrology and urban hydrology. C.E. & Min. E., Natl. Univ. of Colombia, 1949; M.Sc. Rensselaer Polytechnic Institute, 1950; D. Eng. Sc. in Hydromechanics, Columbia, 1955; Purdue Univ. faculty since 1955; Member AGU, ASCE, AWRA, AAAS, IAH, Sigma Xi. Tau Beta Pi. Has been chairman of Fluid Dynamics Committee of Engineering Mechanics Division ASCE and Chairman Task Force on Hydraulics of Bridges ASCE; is Member of Urban Water Resources Research Council ASCE; Member US National Committee for the IASH; 81 journal or conference proceedings publications, 9 published by AGU; 51 technical reports. Coauthor textbook on modeling hydrologic time series. Fellow ASCE. Served as Member of AGU Surface Hydrology Committee since 1975; Chairman of Urban Hydrology Section Executive Committee; coauthor of AGU Water Resources Monograph 7 on Urban Storm-water Hydrology; Convener and chairman of several symposia in Urban Hydrology and on stochastic hydrology.



"At present, the Hydrology Section of AGU is under able direction. Should I be elected President of the Section, I intend to continue the efforts for a stronger participation in the Union affairs and to ensure the quality of the publications and scientific programs in our conferences and symposia."

"I would like to work with the editors of our research journal, *Water Resources Research* (WRR), to accelerate the review process and to reduce the time between the submission of the papers and their publication while maintaining the strong reputation of this journal. At the same time, I would like to provide an outlet for publication of quality original articles in applied hydrology. This would bridge the gap between the research hydrologists who publish in WRR and the large number of practicing hydrologists who at present lack an appropriate publication medium within the section."

"Stronger lines of communication between the president and Section members should be established. This could be accomplished by keeping the members informed about the issues through short articles in *Eos*. Advanced publication of agenda of the meetings of the executive committee would foster the opportunity for the Hydrology Section to reflect on their positions on important matters. The members would have a better opportunity to make their views known and to participate in the decision-making process of their Section. Also more opportunities for membership communication should be given at the luncheon meetings."

tunity of the Section members to reflect on their positions on important matters. The members would have a better opportunity to make their views known and to participate in the decision-making process of their Section. Also more opportunities for membership communication should be given at the luncheon meetings."

"Only active committees should exist and they should have well-defined objectives such as organizing scientific sessions, symposia, Chapman conferences, monograph publications, etc. Consideration should be given to the merging of committees with overlapping objectives while encouraging the formation of interdisciplinary groups, or groups exploring new developments."

"Closer interaction with other sections of the Union should be developed, primarily with the Atmospheric Sciences and Ocean Sciences sections. Hydrologists can contribute much to the study of the coupling of atmosphere-land-ocean systems. This interaction could take the form of joint scientific sessions, joint articles in *Eos*, and invited reviews or editorials in WRR."

"Finally, the presidency of the section cannot be considered as an honorary position but rather as a challenge."

Marshall E. Mass A member of AGU since 1968; 43 years old; Chief of Surface Water Branch, U.S. Geological Survey (USGS); Water Resources Division; design of data networks and mixtures of stochastic and deterministic hydrologic models. B.S. in civil engineering, Clemson, 1963; M.S. in civil engineering, Arizona, 1969; Ph.D. in civil engineering, Colorado State, 1973. Hydrologist with USGS from 1973 to present. Member: AGU, ASCE, AAAS, ASA, Sigma Xi. Has served on Hydrology Committee of AMS, as Rapporteur for Network Design for WMO, Co-director of UNESCO Symposium on Droughts in Arid Lands and currently serves as a member of the U.S. Committee for IAH, Rapporteur for Technical Regulations for WMO Commission on Hydrology, U.S. delegate to WMO Congress, head of U.S. delegation to ISO Technical Committee 113, hydrology in open channels, 48 publications, 9 published by AGU; recently published by WMO, "Concepts and Techniques in Hydrological Network Design"; by USGS, "Design of Surface-Water Data Networks for Regional Information." Served as Chairman and Member of AGU Tellers Committee, Hydrology Section's Network Design Committee and Horton Award Committee; as member of AGU Horton Medal Committee and Hydrology Section's Surface Water Committee and Stochastic Hydrology Work Group; was convener of Chapman Conference on Design of Hydrological Data Networks and served as co-editor of the Conference Proceedings published in WRR.

"The Symposium on Climate and Paleoclimate of Lakes, Rivers, and Glaciers will be held June 4-7, 1984, at Igls, Austria, near Innsbruck. Abstracts of papers to be presented at the meeting should be submitted by December 15 and should deal with any of the following symposium topics related to lakes and rivers or glaciers and ice sheets: present climate, water/ice and energy budgets; recent changes; historical changes; late glacial and post-glacial development; or the physical bases and models for paleoclimate reconstructions and political considerations."

The International Association of Meteorology and Atmospheric Physics International Commission on Climate is sponsoring the event in cooperation with the International Union for Quaternary Research Paleoclimate Commission, the International Association for Hydrological Sciences Commission on AGU Water Resources Monograph Series. He is a member of AWRA and has served on numerous committees for AGU, AWRA, and ASCE. He is now the incumbent Secretary of the Hydrology Section.

Edward A. McBean A member of AGU since 1968; 37 years old; Professor of Civil Engineering, Univ. of Waterloo, Ontario. Major interests: modeling of water resources systems (quantity and quality), cost-benefit assessments. B.A.Sc. in civil engineering, British Columbia, 1968; S.M. in 1970, and Ph.D. in 1973, in civil engineering, MIT. Engineer with Acres Consulting Services 1970-1971; project engineer with Meta Systems, Inc., to 1974; research associate, Water Resources and Marine Sciences Center, Cornell Univ., 1974; Univ. of Waterloo since 1974. Edited four books; 52 refereed papers, 3 published by AGU. Member, organizing committee for International Symposium on Risk and Reliability in Water Resources, Waterloo, June 1978; International Association of Water Pollution Research, Toronto, June 1980; and International Symposium on Real-Time Operation of Hydrosystems, Waterloo, June 1981. Appointed representative to Environment Canada Technical Workshop on Modelling, 1976; and Technical Workshop on Streamflow Forecasting, 1983. Reviewer of technical papers for AGU (from 1974), ASCE, WASP, GWRM, CJCE, CWRA, and WPRC. Associate Editor, *Water Resources Research*, 1979 to present.

"The second reason is the strength of our publication, *Water Resources Research*. Prior to the initiation of WRR, there was no widely distributed publication treating scientific hydrology per se. Articles could be found scattered among engineering journals, conference proceedings, and even in some issues of AGU's *Journal of Geophysical Research*. To be moderately well read in hydrologic literature, one had to diligently search through many journals. However, with the very first issue of WRR, a single source where one could encounter the latest knowledge of all phases of hydrology and its applications in water-resources management was realized. Thus, AGU's *Water Resources Research* is where the science is."

"However, in recent years, each of these two facets of the Hydrology Section have ceased to be wholly true. There are now entities that offer hydrologists affiliations that approach the level of professional value that once was the Hydrology Section's alone. There are also journals dedicated to hydrology and water resources that have made significant inroads into the publication realm that only a few years ago was almost wholly that of WRR."

"Is this unique status of the Hydrology Section to be maintained? I don't know the long-term answer to this question; however, I don't see any entity to which I want to trans-

Meetings

Mantle Processes

A Penrose Conference on "Processes and Evolution in the Mantle" will be held April 29-May 4, 1984, at Gold Canyon Ranch in central Arizona. Attendance is limited to 80, and the deadline for submitting applications is October 31.

The conference is cosponsored by the Geological Society of America and the U.S. Geological Survey (USGS) to bring together petrologists, experimentalists, geophysicists, geochemists, and other researchers to discuss current research of mantle heterogeneities and the processes that create them. On the agenda is a one-day field trip to the site of the San Carlos, Ariz., mantle xenolith.

Proposed sessions will address the petrology and geochemistry of mantle material that has been modified by multistage events, geochemical evidence of multistage mantle processes from the study of lavas, experimentally determined evidence of the geochemistry of mantle metasomatic fluids, geophysics of melt processes and melt migration, and evolution of chemical heterogeneities in a convecting mantle.

Applications, along with a brief description of topics to be presented at the conference or reasons for attending, should be sent to Jane E. Pike, U.S. Geological Survey, 345 Middlefield Rd., MS 75, Menlo Park, CA 94025.

Deadline for nominations

is November 15, 1983.

Physics of Magna Transfer Wednesday

P.M., Cathedral Hill (cosponsored by T.V.)

Electrical Properties of the Crust and Mantle Friday A.M., Holiday Inn (cosponsored by G.P.)

Physical Chemistry of Minerals Friday P.M., Holiday Inn (cosponsored by T.V.)

Nominations for the James B. Macelwane Award

Nominations for the prestigious James B. Macelwane Award are still being accepted. Up to three recipients may receive this honor per year to assure that individuals in all areas of geophysics are regularly recognized.

The Macelwane award is given by AGU in "recognition of significant contributions to the geophysical sciences by a young scientist of outstanding ability." Recipients must be less than 36 years old.

Letters of nomination, outlining significant contributions and curriculum vitae should be sent directly to:

J. Freeman Gilburt
IGPP A-025
University of California, San Diego
La Jolla, CA 92093

Atmospheric Sciences

Lightning: The Workman Memorial Session Monday P.M., Cathedral Hill, Eldorado Room This special poster session on lightning is being held to honor the memory of E. J. Workman, 1899-1982, who was one of the early leaders of modern research on thunderstorms, cloud physics, and atmospheric electricity. Much of the recent progress in lightning and thunderstorms research stems from the work of Workman and his colleagues S. E. Reynolds and I. Langmuir. New Mexico Institute of Mining and Technology over the last 4 decades. Their legacy includes the establishment of Langmuir Laboratory at Socorro, N. Mex.

Parallel poster sessions (Friday A.M., Cathedral Hill, El Dorado Room) will feature Charged-Particle Populations and Precipitation; Waves and Instabilities in Laboratory and Space Experiments; Auroral Hiss and Kilometric Radar; and Geomagnetic Pulses (37 papers in all). Bob Spilo will chair.

Parallel poster sessions (Friday P.M., Cathedral Hill, International Room) will cover the Geomagnetic Tail and Boundary Layer; Ionospheric Radar Results, and Waves and Instabilities in the Ionosphere and Magnetosphere (52 papers in all). Umran Inan will chair.

Tectonophysics

Physics of Magma Transfer Wednesday P.M., Cathedral Hill, International Room (co-sponsored by the Volcanology, Geochemistry, and Petrology and Tectonophysics sections).

This symposium will cover the physics of magma transfer from the atomic scale to the scale of mantle plumes. Topics range from experiments to fluid mechanical models of diapir and

plume with Space Plasma, John Olson (Pc 1) and Gordon Rostoker (Pc 4-5 and P. 21) are invited speakers; Joseph Kan, who organized the session, will chair.

ISEE-3 Observations of the Distant Magentotail (Friday A.M. and P.M., Cathedral Hill, Japanese Pavilion) will highlight the final day of the meeting. The two oral sessions at geocentric distances of 80 and 240 earth radii in the tail. Invited speakers are Tsvy Oya (Planet A and MST-5), R. Reinhard (Giotto), T. von Rosenvinge, Bruce Tsurutani, M. A. Green, Ed Jones, Fred Scarf, P. W. Daly, and George Glocer. Eleven contributed papers are interspersed in the program. Papers and posters will be presented in the program.

George Parks and von Rosenvinge will chair.

Parallel poster sessions (Friday A.M., Cathedral Hill, International Room) will feature the Geomagnetic Tail and Boundary Layer; Ionospheric Radar Results, and Waves and Instabilities in the Ionosphere and Magnetosphere (52 papers in all). Umran Inan will chair.

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GAP

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Exploration Geophysics

GEOPHYSICAL METHODS
RADIATION IMPEDANCE OF VERTICALLY VIBRATING SEISMIC SOURCES
Jeffrey A. Dunn (Geophysical Engineering Geoscience, University of California, Berkeley); presently Exploration and Production Research, ARCO Oil and Gas Company, P.O. Box 2019, Dallas, TX 75211

The thickness and propagation velocity of a surface layer can theoretically be determined from seismic radiation impedance measurements using a vertical vibrator. These studies also provide physical insight into vibrator-earth interaction.

The radiation impedance of a circular disc vibrating vertically on an elastic half-space is calculated with a spacing that is a function of the ratio between half-space radius and seismic wavelength. An impedance function is nearly independent of the spacing; therefore, the magnitude is affected. At high frequencies the impedance depends strongly on the flexibility of the base, and at low frequencies the base introduces an additional resonant effect, the frequency of which is a function of the baseplate mass.

The presence of a surface layer produces an impedance contrast. The amplitude of the oscillations is a function of the acoustic impedance contrast and depends upon the radiation pattern of the source. The oscillations are excited by a vertical vibrator, which has a vertical base and both the period and amplitude of the oscillations are inversely proportional to the base radius. The amplitude and frequency determine the spacing required to obtain resonance.

With impedance measurements over a sufficiently broad frequency range (for about one octave), it can be determined what type of oscillations and the layer thickness to determine the shear velocity and thickness of the layer of material beneath the baseplate.

Geophysical Research Letters

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Climatic Effects of the Eruption of El Chichon

The El Chichon Volcanic Cloud: An Introduction (Paper 3L1457)
James B. Pollack, David B. Trivelpiece, David J. Hofmann, and James M. Rosen

Formation of the El Chichon Aerosol Cloud (Paper 3L1491)
C. A. Barth, R. M. Jakobson, R. J. Thomas, G. E. Thomas, B. M. Jakobson, and R. A. West

Satellite Lidar Scanning the Thermal Infra-red Emissions from the El Chichon Volcanic Cloud (Paper 3L1492)
R. E. Taylor, J. E. Turner, and R. W. Sanders

The El Chichon Cloud Over Central Europe: Observed by Lidar at Garnier/Sarissenches During (Paper 3L1493)
O. Reiter, H. Jäger, W. Cornforth, and W. Frank

One-Year Lidar Observations of the Stratospheric Aerosol Cloud at Frascati, March 1982-March 1983 (Paper 3L1572)
A. Adamek, F. Cimarelli, G. Flores, and G. P. Webb

Spectroscopic Measurements of the 8- to 13-Micrometer Transmission of the Upper Atmosphere Following the El Chichon Eruption (Paper 3L1497)
C. W. Wilson, E. D. Blackford, and J. H. Mum

Spectral Extinction of Direct Solar Radiation by the El Chichon Cloud During December 1982 (Paper 3L1498)
J. E. Taylor, D. L. Hart, and J. D. Pollard

Stratospheric Aerosol Light Absorption Before and After El Chichon (Paper 3L1571)
A. D. Clarke, R. J. Charlton, and J. A. Ogren

Effect of the Eruption of El Chichon on Stratospheric Aerosol Size and Composition (Paper 3L1588)
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Measurements of the Aerosol Size Distributions in the El Chichon Cloud (Paper 3L1589)
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Changes in the Sub-2.5 Micron Diameter Aerosol Observed at 20 km Altitude After the Eruption of El Chichon (Paper 3L1590)
El Chichon Volcanic Ash in the Stratosphere: Particle Abundance and Size Distributions After the 1982 Eruption (Paper 3L1591)
B. W. Galvin, M. A. Kritz, and A. L. Lazarus

Balloon and Aircraft Measurements of Stratospheric Sulphur Mixture Ratio Following the El Chichon Eruption (Paper 3L1592)
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Size-Specific Composition of Aerosols in the El Chichon Cloud During December 1982 (Paper 3L1593)
James F. Verdin, Estelle P. Coulam, Edward C. Yiu, Kevin D. Taylor, and Mark A. Kritz

Measurements of Stratospheric SO₂ After the El Chichon Eruptions (Paper 3L1594)
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A Two-Dimensional Model Simulation of the El Chichon Volcanic Cloud in the Radiation Budget of the Northern Tropics (Paper 3L1596)
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Possible Effects of the El Chichon Volcanic Cloud on the Radiation Budget of the Northern Tropics (Paper 3L1597)
John P. Wallace, James R. Pollack, and Thomas P. Ackerman

Regular Issues

No Water, No Granules—No Continuum (Paper 3L1622)
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Post-Glaciogenic Results from the Jurassic Topley Intrusive in the Sukkinia Subterrane of Belarus (Paper 3L1620)
David T. A. Synnott